

The Case for a Super Neutrino Beam

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Frascati Physics Series

HEAVY QUARKS AND LEPTONS – San Juan, Puerto Rico, June 1 - June 5, 2004

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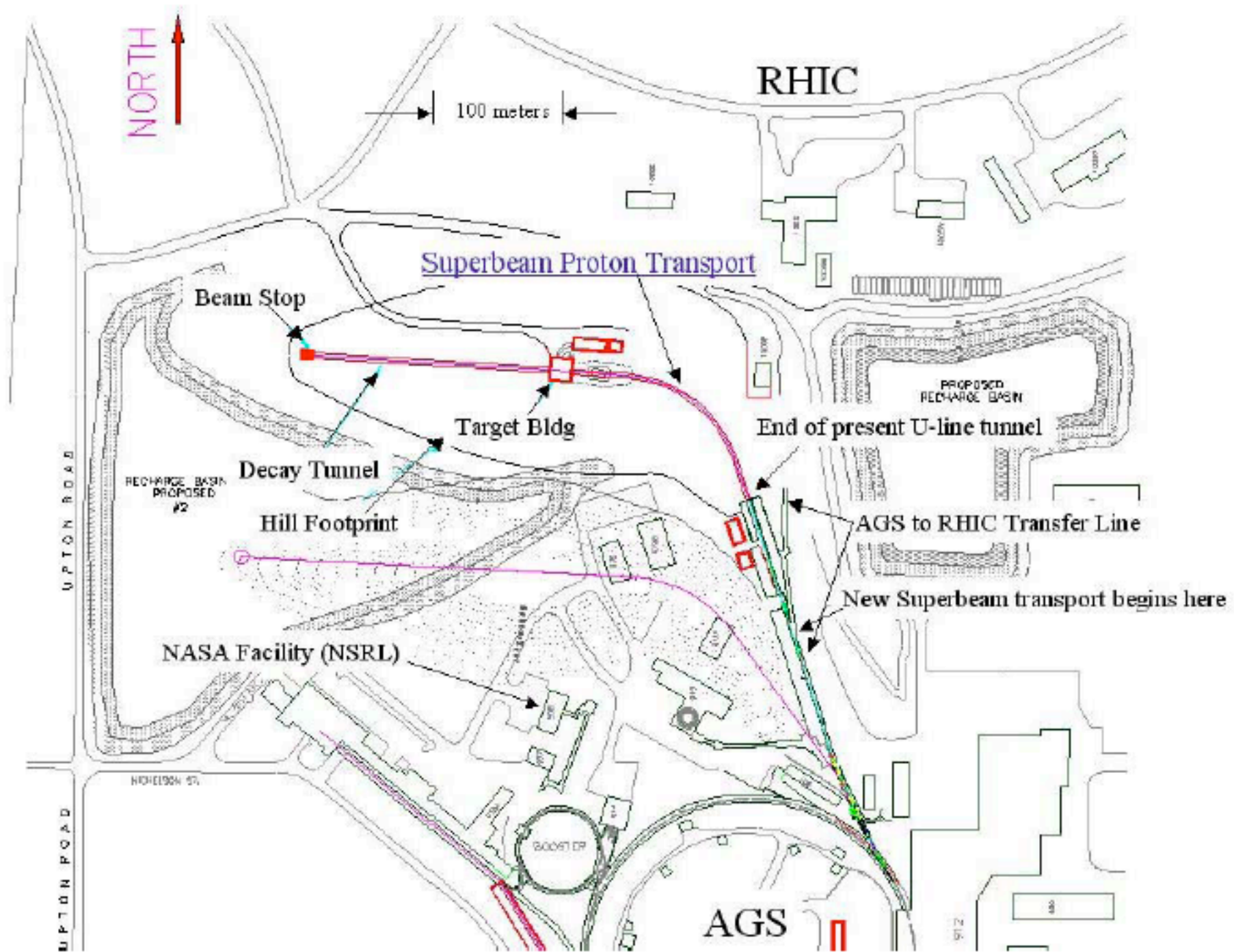
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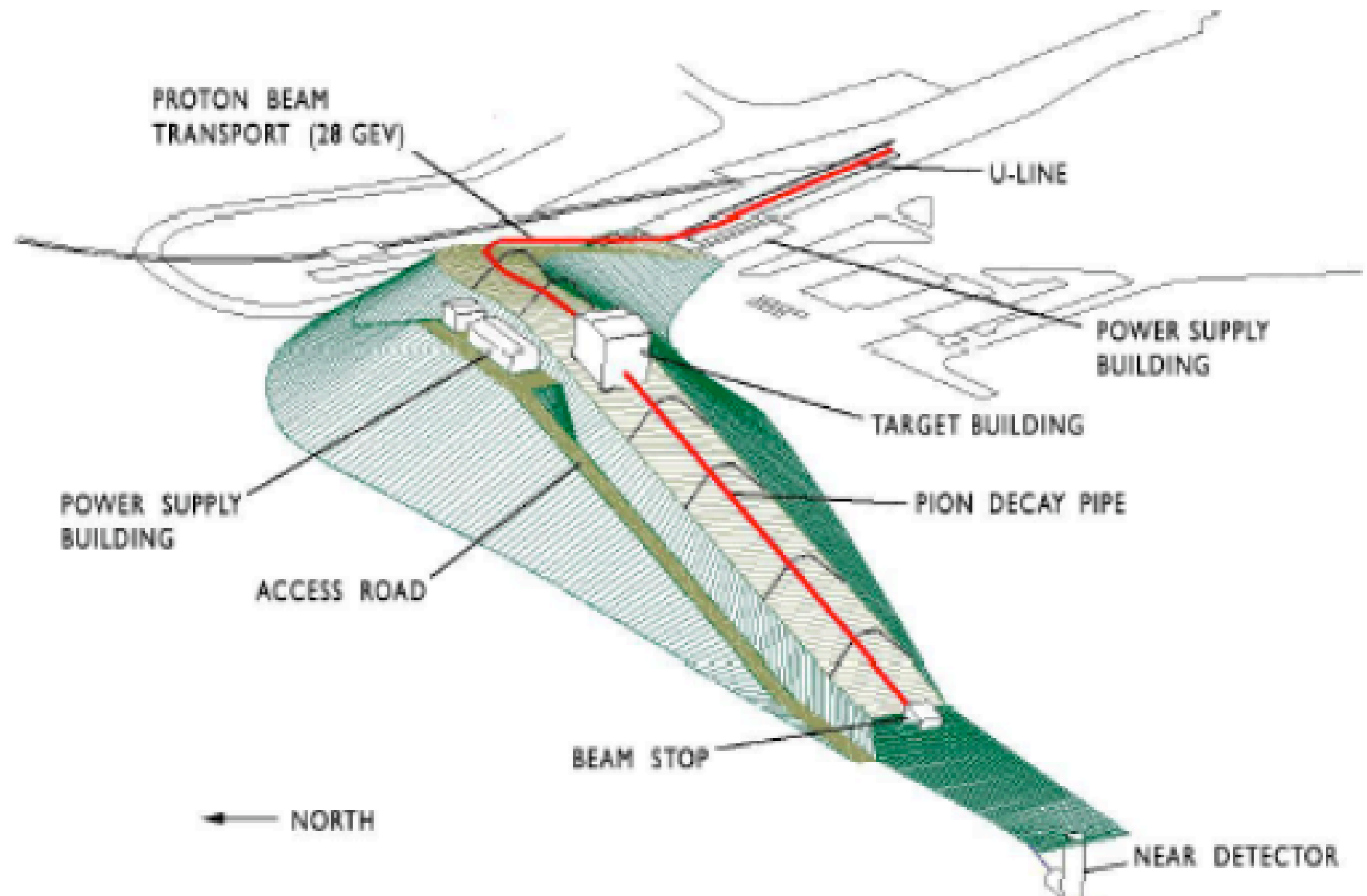
ABSTRACT

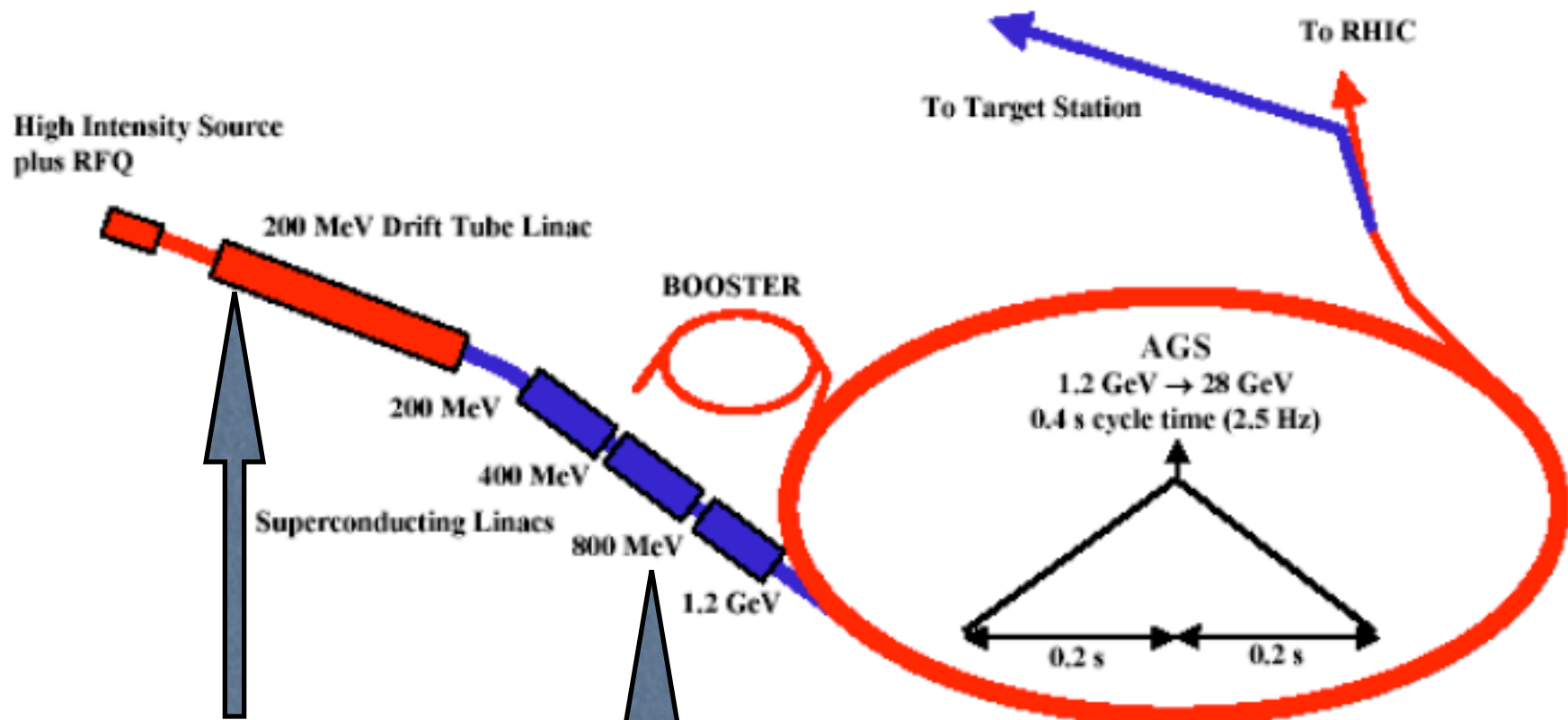
In this paper I will discuss how an intense beam of high energy neutrinos produced with conventional technology could be used to further our understanding of neutrino masses and mixings. I will describe the possibility of building such a beam at existing U.S. laboratories. Such a project couples naturally to a large (> 100 kT) multipurpose detector in a new deep underground laboratory. I will discuss the requirements for such a detector. Since the number of sites for both an accelerator laboratory and a deep laboratory are limited, I will discuss how the choice of baseline affects the physics sensitivities, the practical issues of beam construction, and event rates.

Update on AGS based Super Neutrino beam

- New conceptual design document
BNL-73210-2004-IR. (sent to DOE)
- http://raparia.sns.bnl.gov/nwd_ad
- Redesigned beam facility: more compact,
now possible to make decay pipe longer.
- Completely new design for injector LINAC:
cheaper and faster to build.







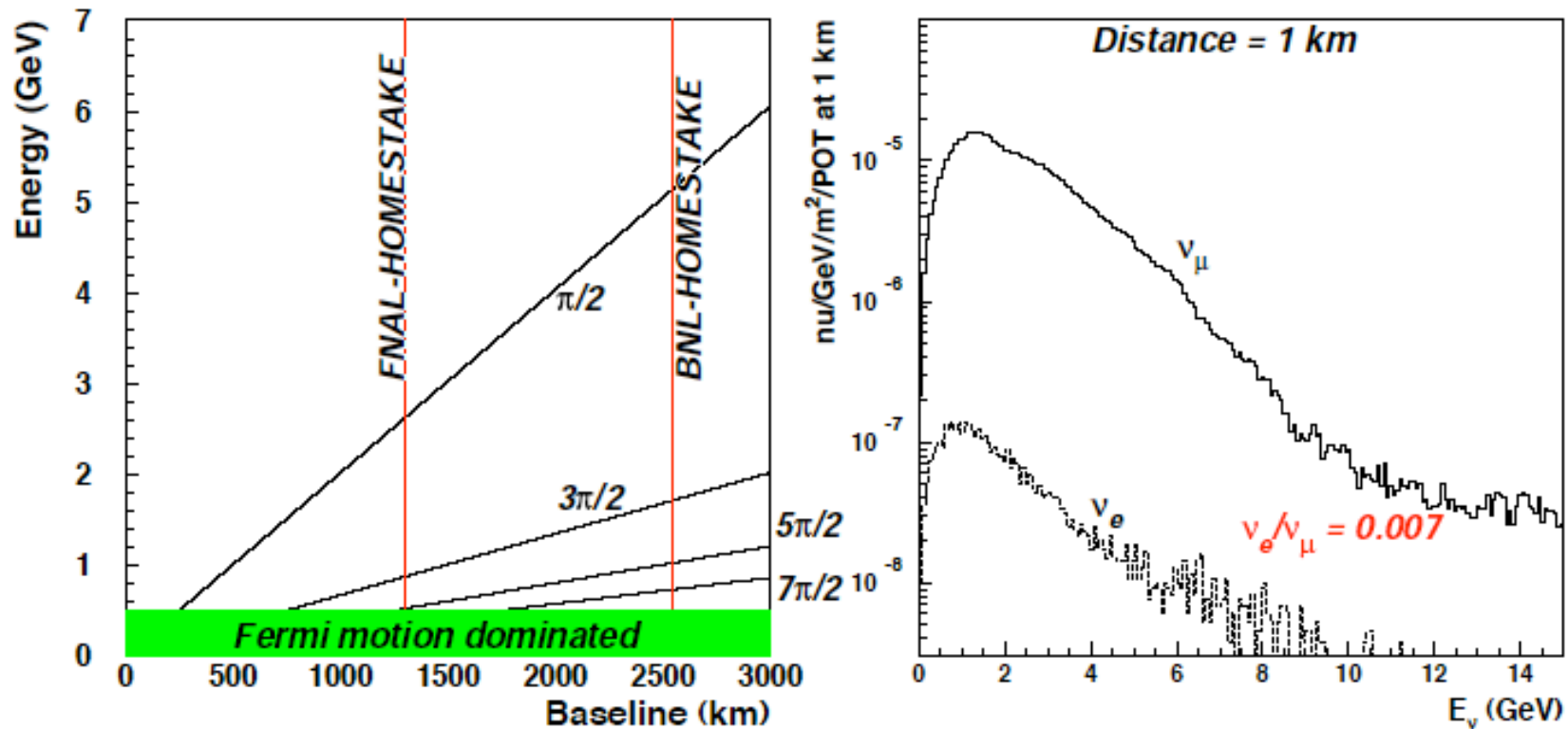
New idea: extend existing
RTL to 400 MeV using a
coupled cavity linac just
like FNAL

Old: low beta: 805 MHz
medium high beta: 1610 MHz

New idea: after 400 MeV use 805 MHz all
the way to ~1.5 GeV.
Use SNS design and get to higher energy

Back to Very Long Baselines

Oscillation Nodes for $\Delta m^2 = 0.0025 \text{ eV}^2$ BNL Wide Band. Proton Energy = 28 GeV



Use same spectrum to study both baselines for this study. Comment on useful spectrum changes. Use 500 kT for both baselines with same performance.

Simple rules

- Multiples nodes important for precision and new physics.
- Long distances separate CP and matter effects.
- Need $2500 \text{ kT} * \text{MW} * (10^7 \text{ sec})$ for measuring CP (regardless of distance and value of θ_{13})
- For CP violation study NO conventional beam experiment can get below $\sin^2 2\theta_{13} \sim 0.01$

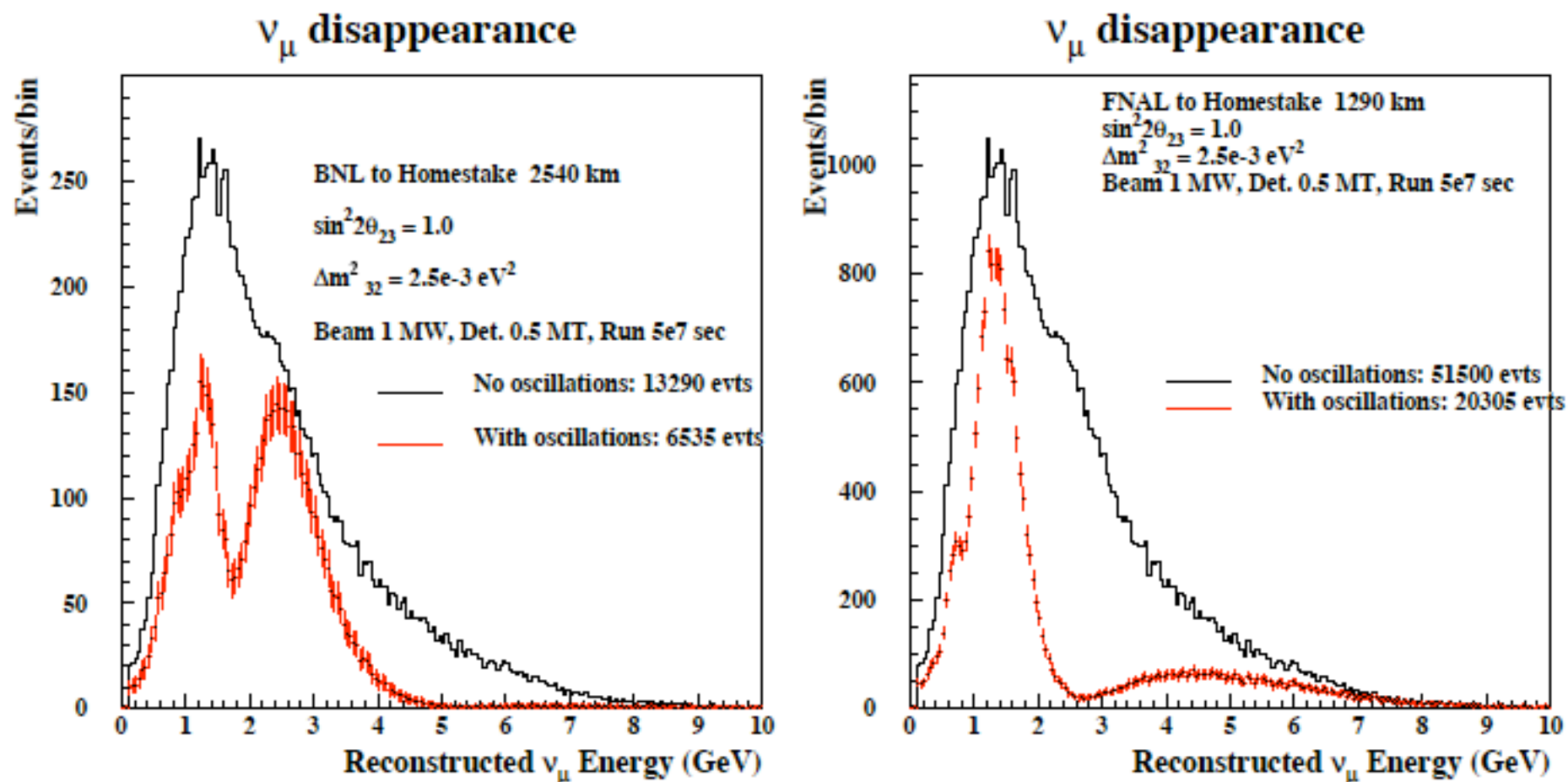


Figure 2: *Simulated spectrum of detected muon neutrinos for 1 MW beam and 500 kT detector exposed for 5×10^7 sec. Left side is for baseline of 2540 km, right side for baseline of 1290 km. The oscillation parameters assumed are shown in the figure. Only clean single muon events are assumed to be used for this measurement (see text).*

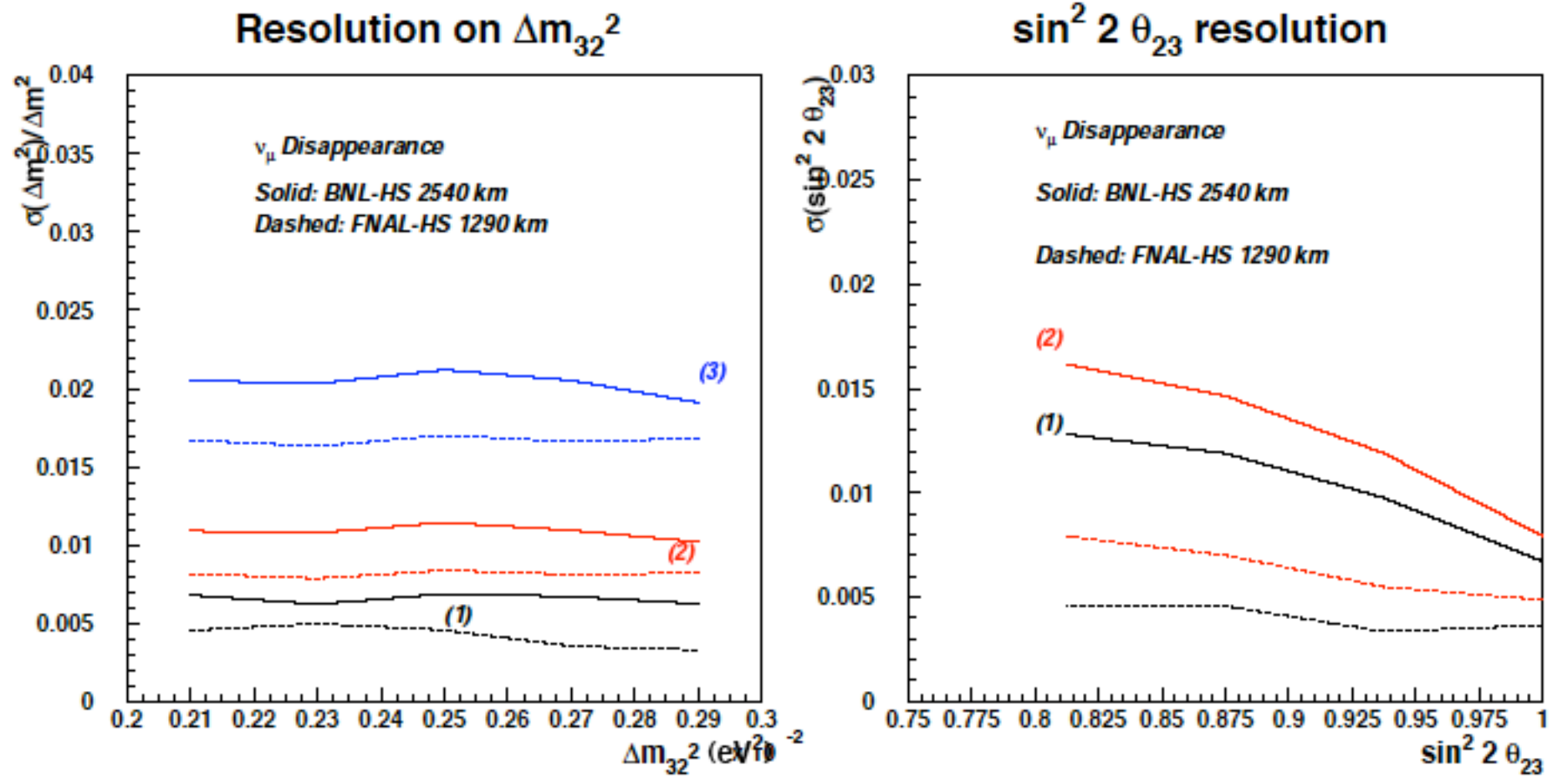
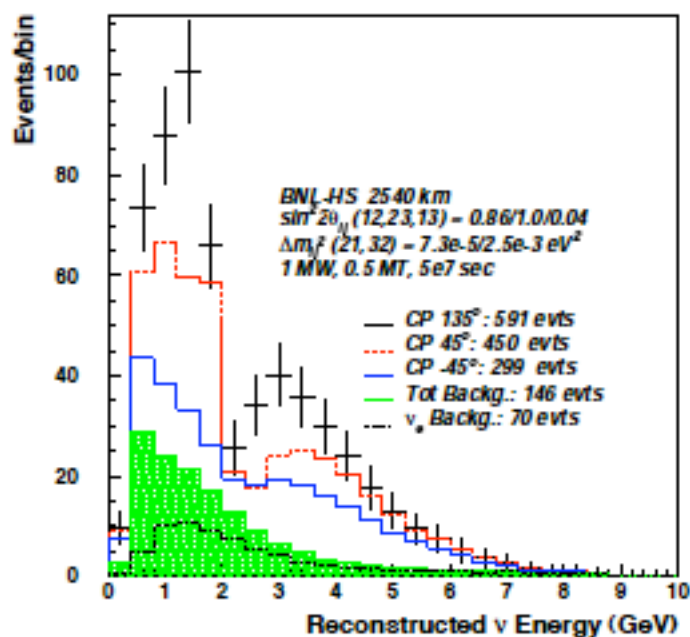
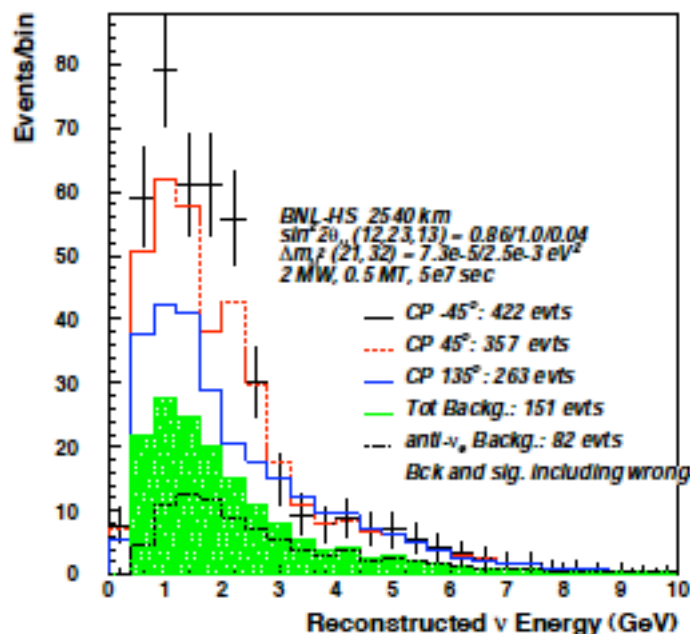


Figure 3: 1 sigma resolutions on Δm_{32}^2 (left) and $\sin^2 2\theta_{23}$ (right) expected after analysis of the oscillation spectra from Figure 2. The solid curves are for BNL-HS 2540 km baseline, and the dashed are for FNAL-HS 1290 km baseline. The curves labeled 1 and 2 correspond to statistics only and statistics and systematics, respectively (similarly for dashed curves of the same color). The curve labeled (3) on the left has an additional contribution of 1% systematic error on the global energy scale.

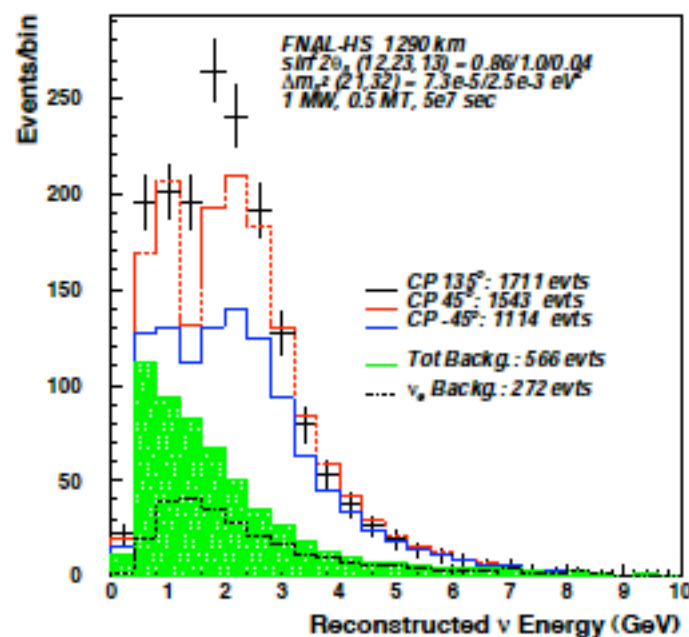
ν_e APPEARANCE



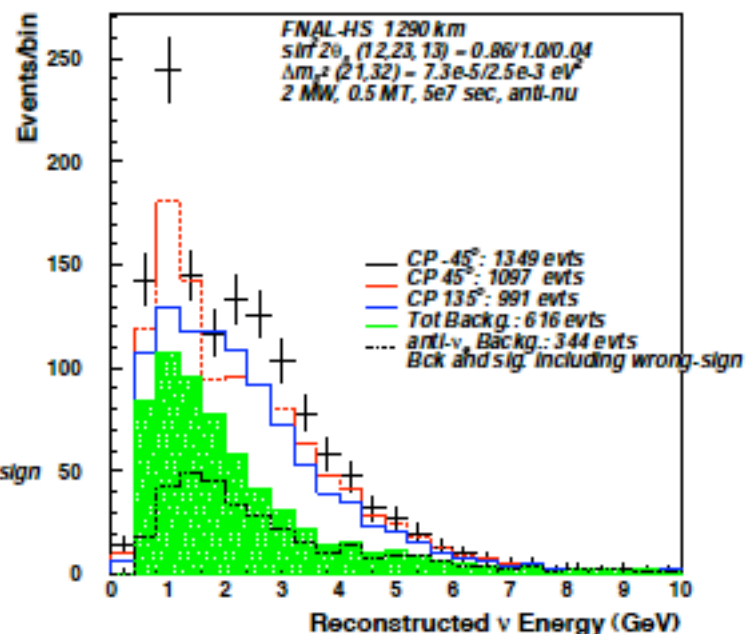
Anti- ν_e APPEARANCE



ν_e APPEARANCE



Anti- ν_e APPEARANCE



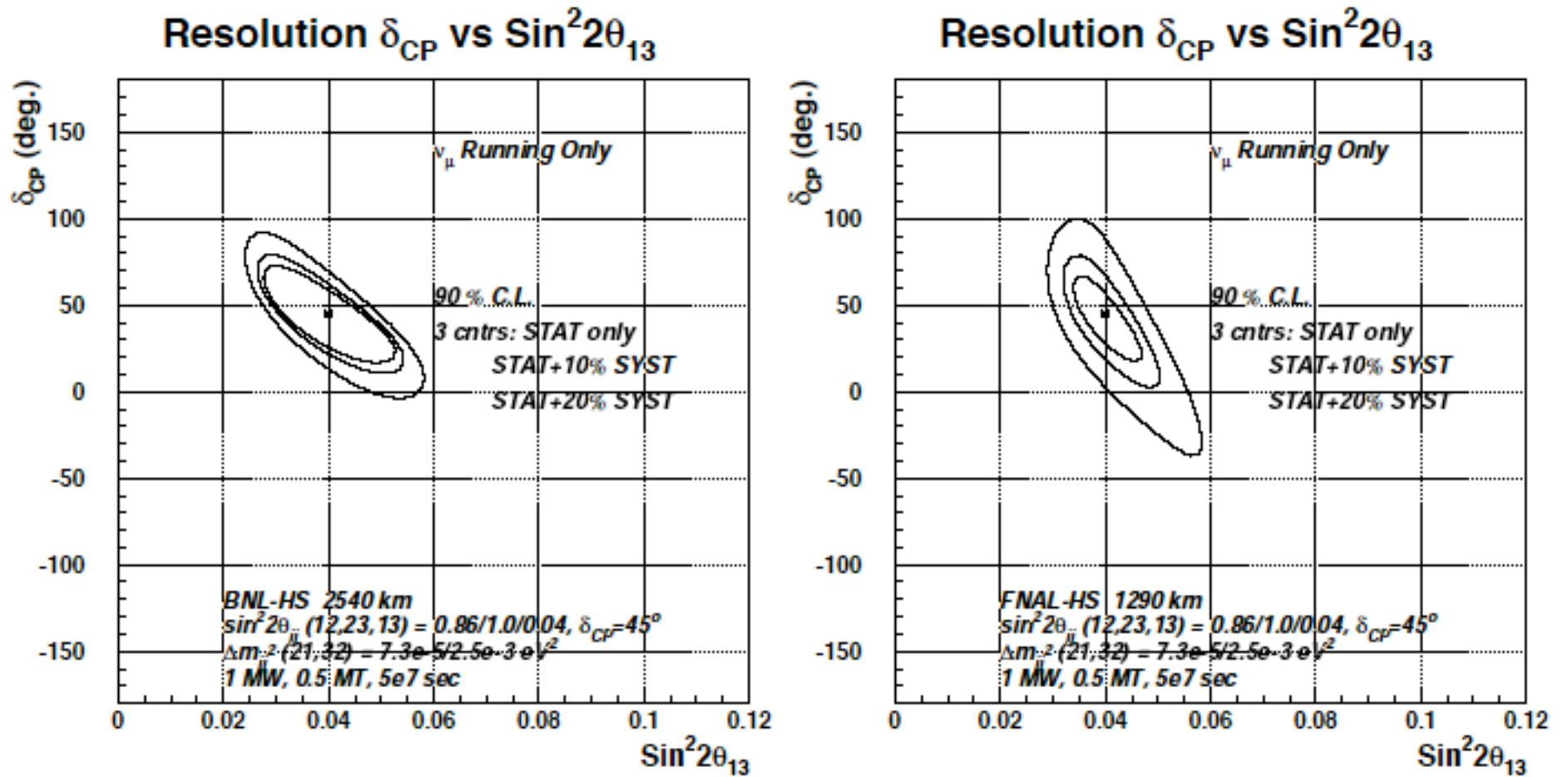


Figure 6: 90% confidence level error contours in $\text{sin}^2 2\theta_{13}$ versus δ_{CP} for statistical and systematic errors with neutrino data alone. Left is for BNL-HS and right is for FNAL-HS. The test point used here is $\text{sin}^2 2\theta_{13} = 0.04$ and $\delta_{CP} = 45^\circ$. $\Delta m^2_{32} = 0.0025 \text{ eV}^2$, and $\Delta m^2_{21} = 7.3 \times 10^{-5} \text{ eV}^2$. The values of $\text{sin}^2 2\theta_{12}$ and $\text{sin}^2 2\theta_{23}$ are set to 0.86, 1.0, respectively.

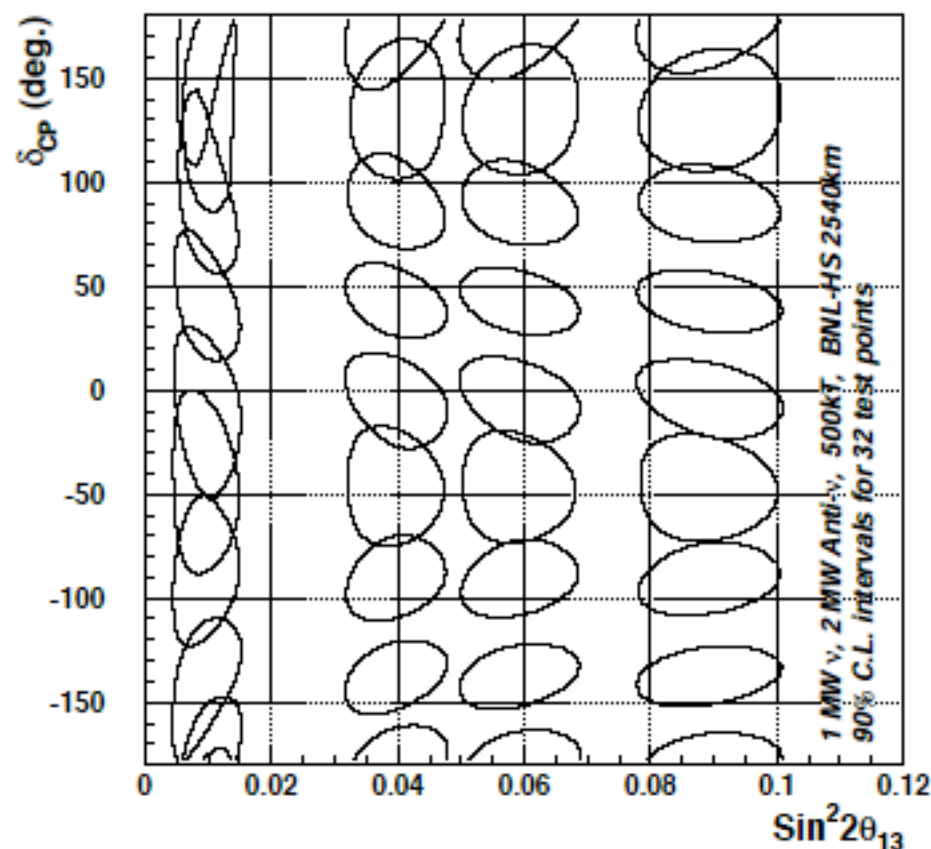
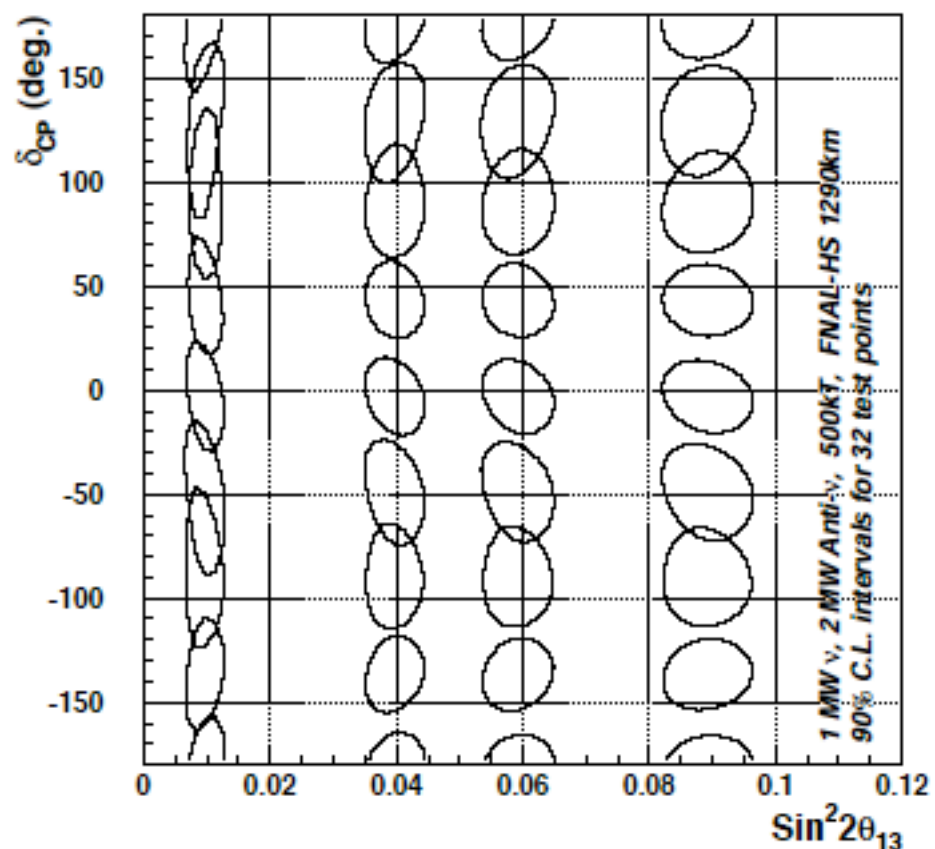
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Figure 7: 90% confidence level error contours in $\sin^2 2\theta_{13}$ versus δ_{CP} for statistical and systematic errors for 32 test points. This simulation is for combining both neutrino and anti-neutrino data. Left is for BNL-HS and right is for FNAL-HS. We assume 10% systematic errors for this plot.

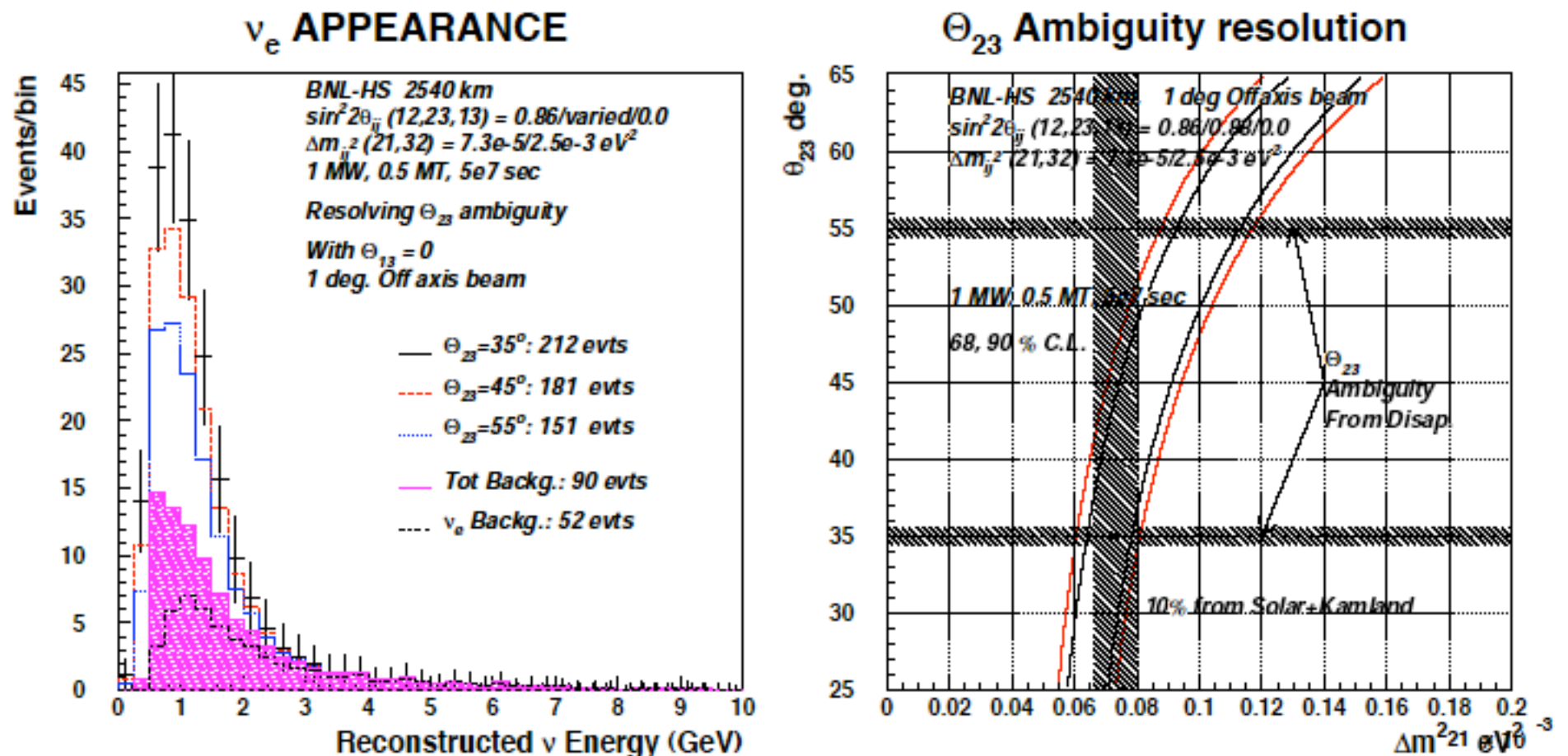


Figure 8: Expected spectrum of electron neutrinos (left) for $\theta_{13} = 0$ and other assumed parameters indicated in the figure. The right hand side shows the resolution of the $\theta_{23} \rightarrow \pi/2 - \theta_{23}$ ambiguity using the measurement of $\sin^2 2\theta_{23}$ from disappearance and assuming a 10% measurement of Δm_{21}^2 from KAMLAND. The area between the curves is allowed by the appearance spectrum (left) for $\theta_{23} = 35^\circ$.

Conclusion

- Shorter baseline somewhat better for measuring mixing angles.
- CP phase measurement statistically indep of distance, but better at longer distance because systematic errors less important.
- With longer distance can get CP phase with neutrino running alone.
- Longer distance better for solar effect and resolving θ_{23} ambiguity.
- Eventually have to couple to DUSEL I MT det.